

From UV to Visible Light: Unveiling the Secrets of Galaxy Size Evolution in the HSC+CLAUDS¹

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Motivation

Evolution of galaxy morphologies tells us about the **physical processes shaping galaxies** in the Universe. Analysis using rest-frames UV (tracing star-forming regions) and optical (tracing the bulk of the stellar population) wavelengths is crucial.

Data ~18 deg² in 6 filters: CFHT CLAUDS² (U) and Subaru HSC³ ($q, r, i \ge 8, y$)

and Subaru HSC³ (g, r, i, z & y) Mass & redshift: M_{*}>10^{9.5}M_o; 0.1<z<0.9 Star-forming galaxies (SFGs): 200,000 Quiescent galaxies (QGs): 80,000 Rest-frames: 3000Å (UV) & 5000Å (optical)

UV (new stars) Optical (old stars)



Figure 1: Sérsic profile fitting using GALFIT⁴ (A) image cut-out, (B) masked image, (C) best-fit models of the target galaxy & bright neighbours and (D) residual image

