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Euclid ERO: a pilot investigation of the HIROs physical properties found in the Perseus field

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H I E R O S



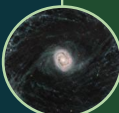
Dust-obscured, star-forming sources at $z \sim 3 - 6$



Undetected by HST: extreme red colors



Detected by Spitzer/IRAC



Limited photometric information: uncertain SED



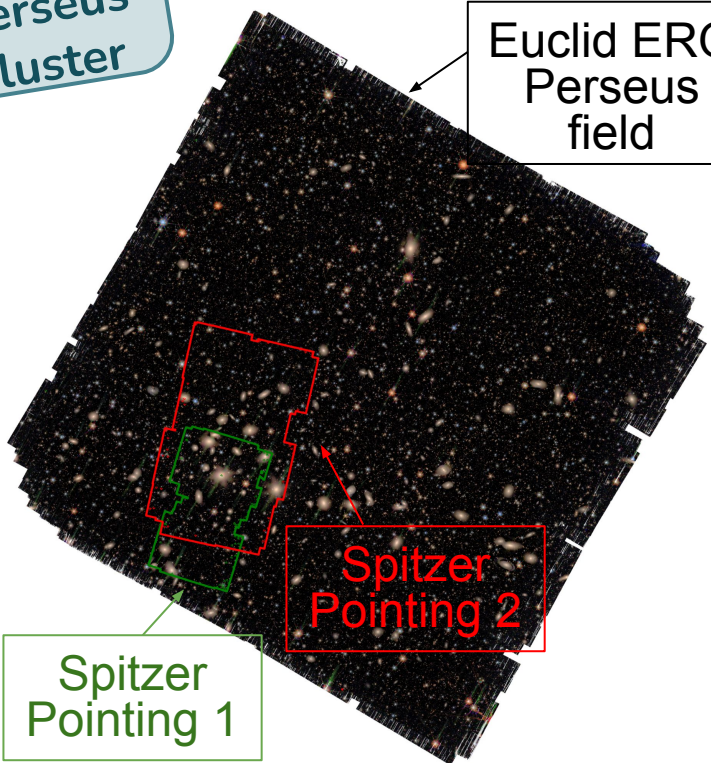
Aim:
statistically
characterizing
this population

Why interesting?
Non-negligible contribution to
the cosmic **SFRD** up to $z \sim 5 - 7$
(Barrufet et al. 2023, Rodighiero et
al. 2023, Wang et al. 2024)

Dominance in the bright end of stellar **MF**
(Rodighiero et al. 2007, Gottumukkala et al. 2024)

Perseus cluster

Euclid ERO Perseus field



2 pointings by
IRAC/Spitzer



Euclid ERO of the
Perseus field
(EC: Cuillandre et al. 2024)

$H - 4.5 \mu\text{m} > 2.25$



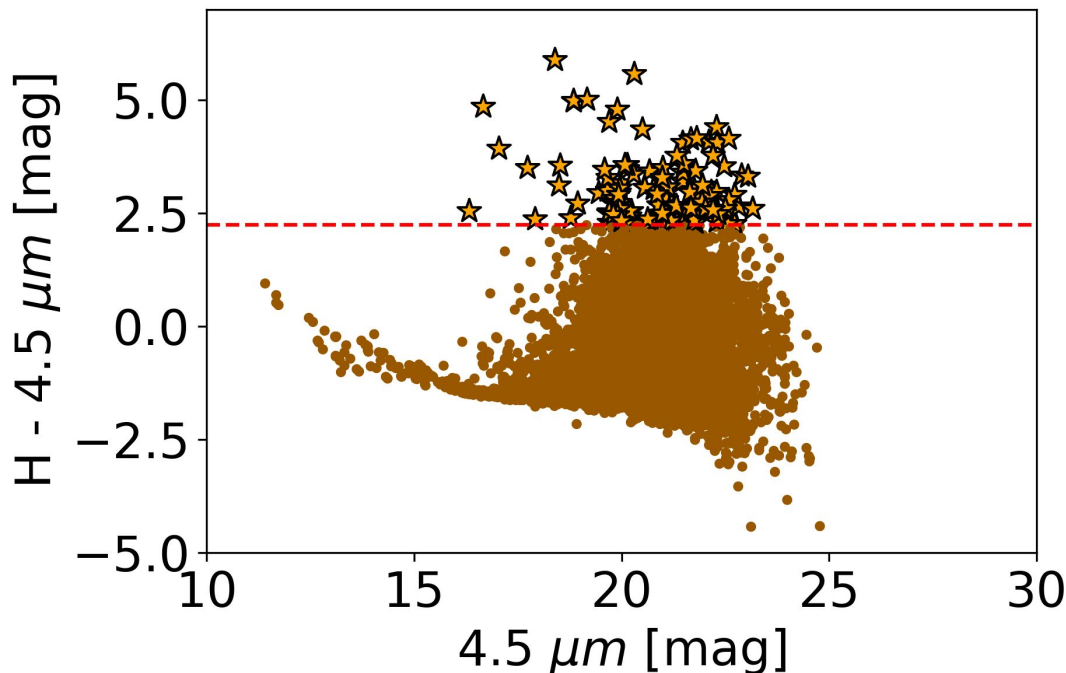
HIEROs

(Wang et al. 2016)

IRAC *Ch2*
magnitude limit:
 ~ 22.6 mag

Euclid preliminary

Color selection



Removing contaminants:

globular clusters,
brown dwarfs



Visual check:

no artifacts, defects

$$-0.1 < VIS - Y < 0.7$$

+

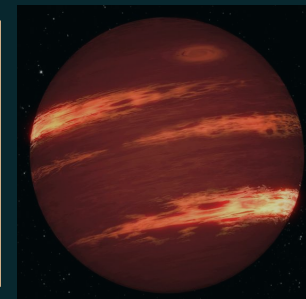
$$-0.3 < J - H < 0.5$$

(EC: Voggel et al. 2024)



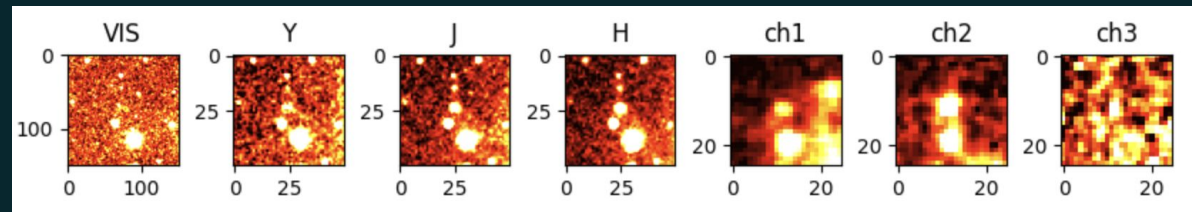
Fit with BD
templates

(Burrows et al. 2006)

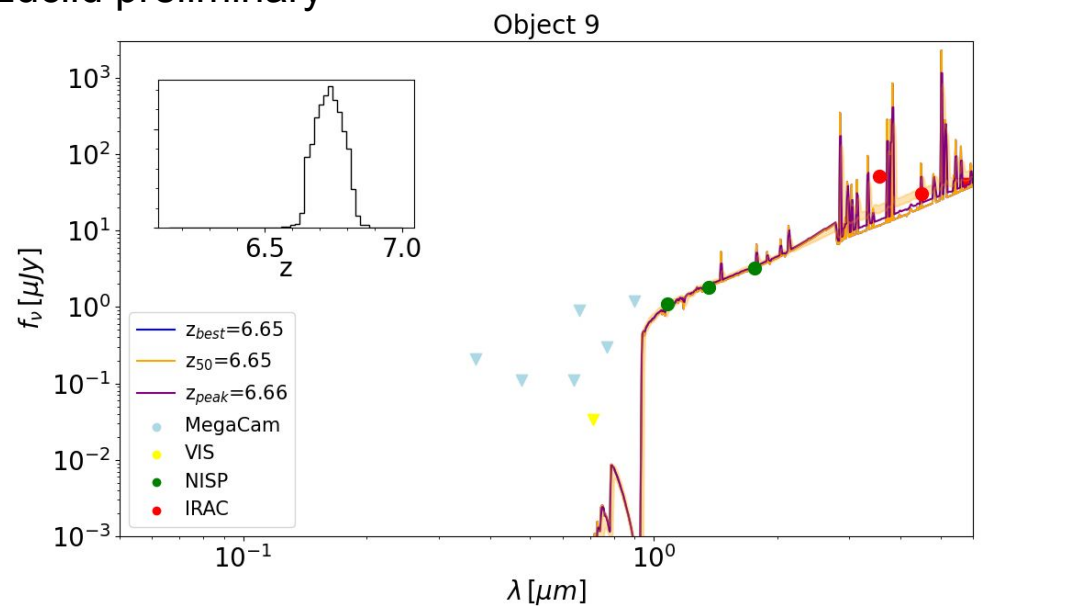


Final sample of **43** objects

Multiwavelength photometry



Euclid preliminary

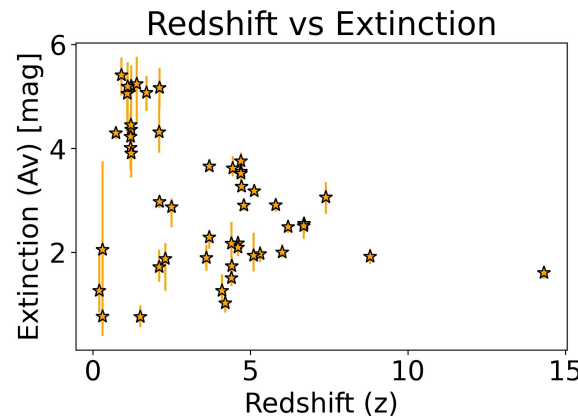


Bagpipes:
SED fitting

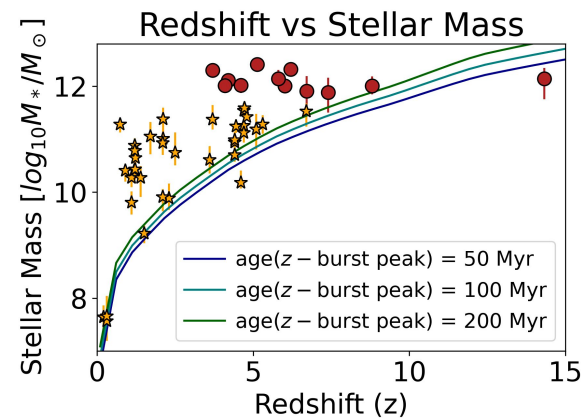
Bagpipes:
results

Massive and dusty galaxies:

mean $A_V \sim 3$ mag
mean $M_* \sim 10^{10.9} M_\odot$



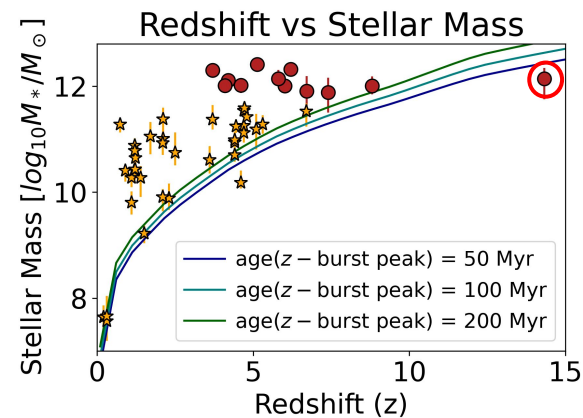
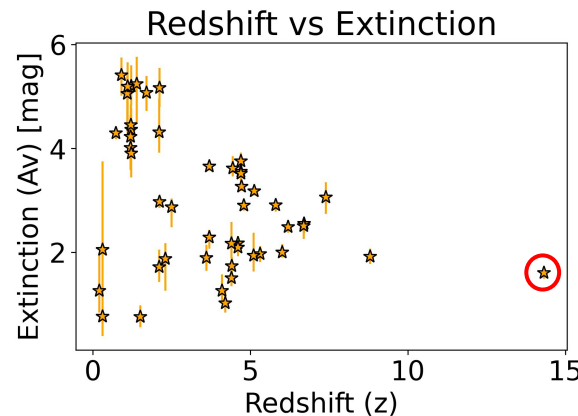
Euclid preliminary



Bagpipes:
results

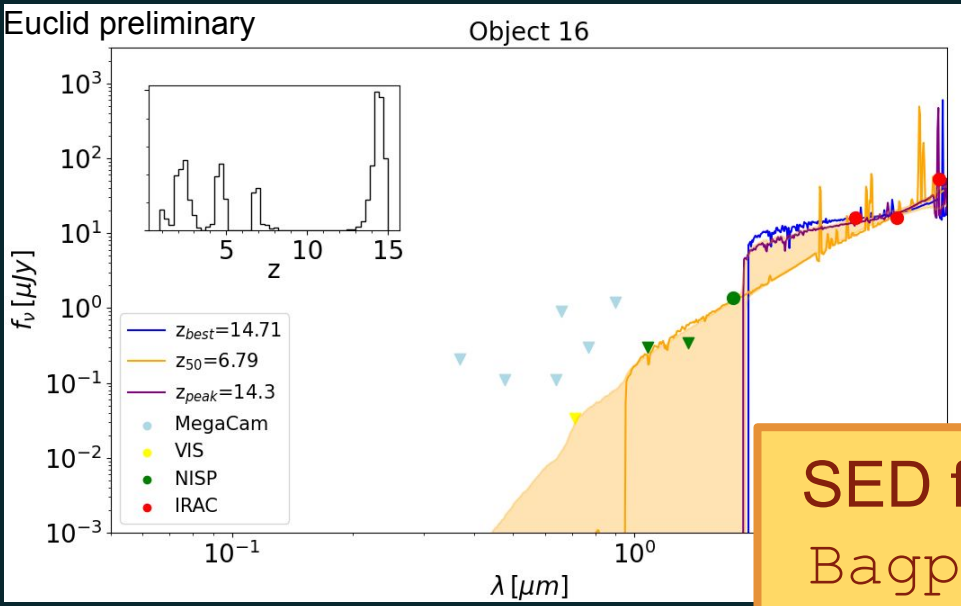
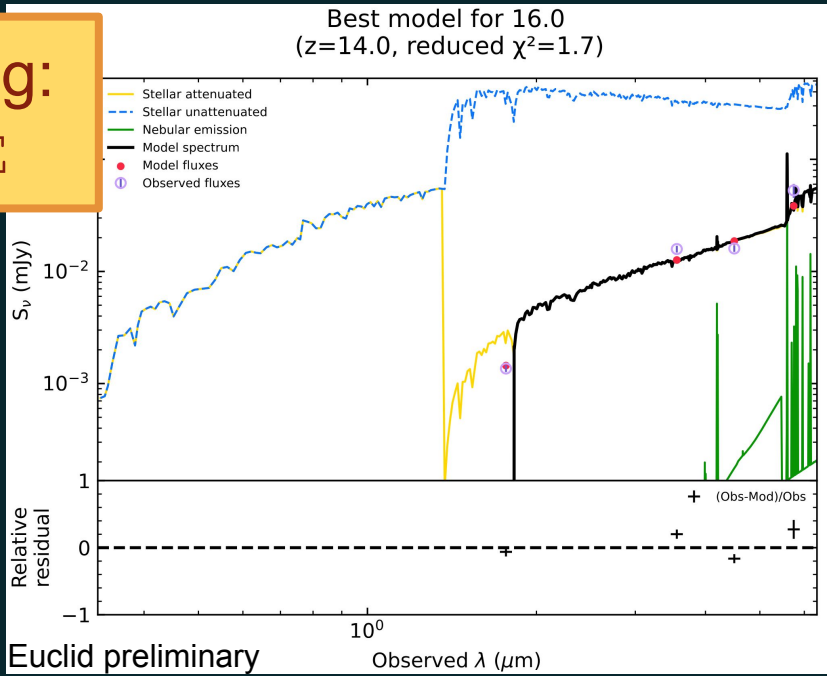
Massive and dusty galaxies:

mean $A_V \sim 3$ mag
mean $M_* \sim 10^{10.9} M_\odot$



Candidate at $z \sim 14$?

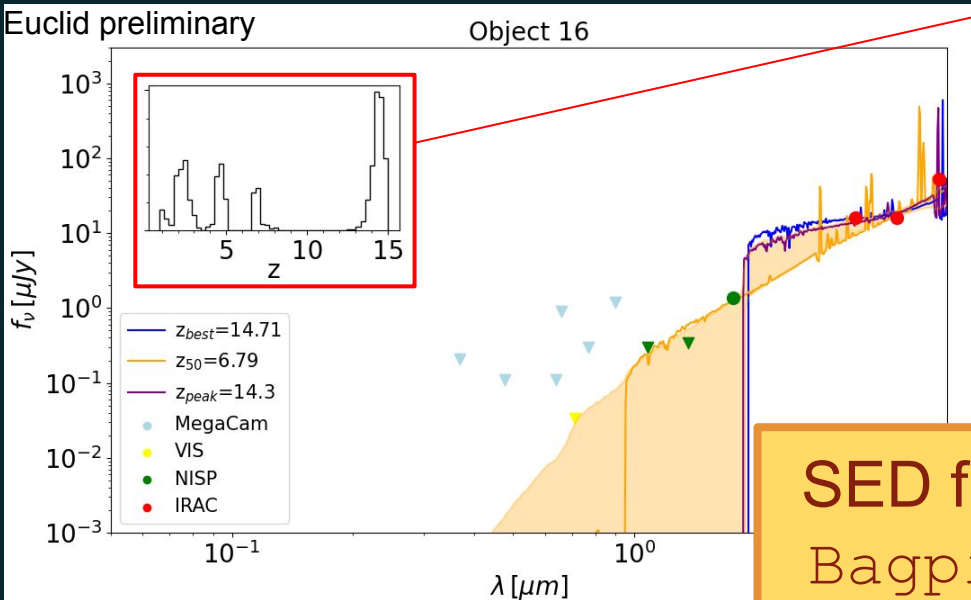
SED fitting: CIGALE



SED fitting: Bagpipes

Candidate at $z \sim 14$?

Multi-peak behaviour: need for spectroscopic confirmation

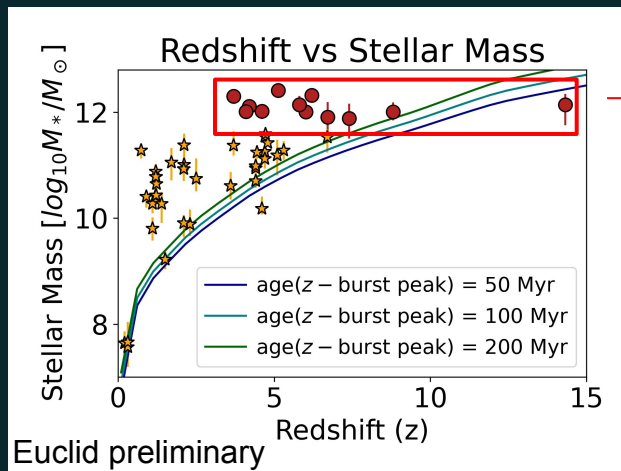


SED fitting:
Bagpipes

Bagpipes:
results

Massive and dusty galaxies:

mean $A_V \sim 3$ mag
mean $M_* \sim 10^{10.9} M_\odot$



Problem:
too high M_*



Problem:
too high M_*



AGN
contribution?



Problem:
too high M_*



AGN
contribution?



SED fitting:
CIGALE with
AGN model
component



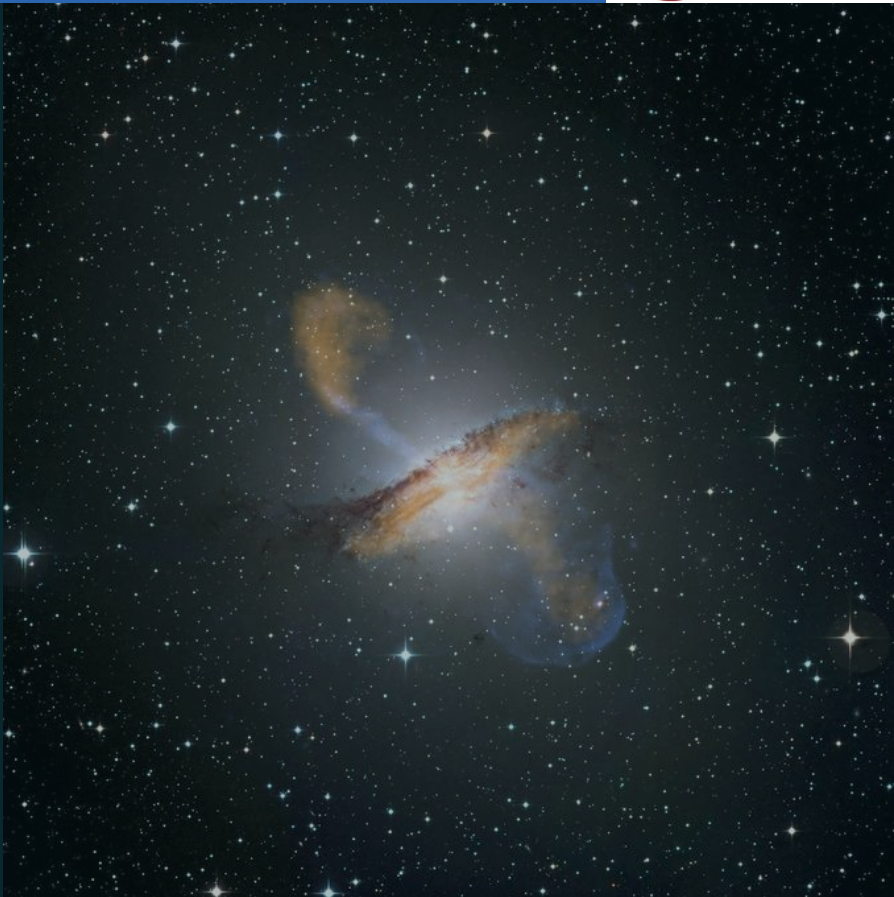
Problem:
too high M_*



AGN
contribution?

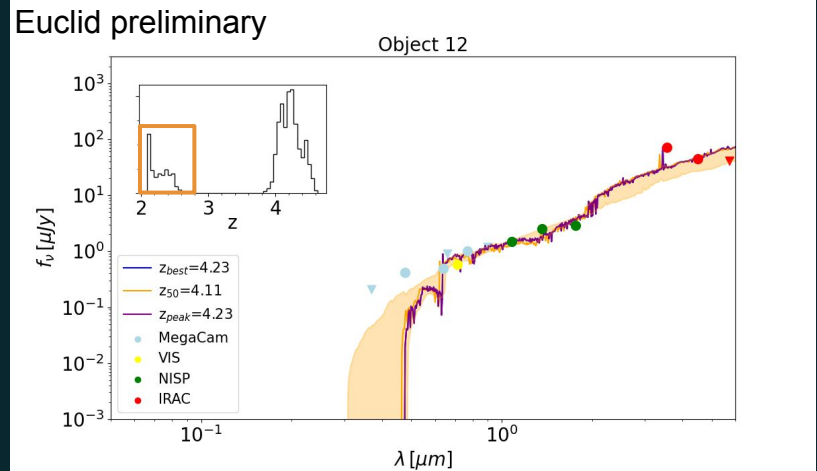


Result:
adding **AGN**
component is
not the solution



Problem:
too high M_*

Solution:
forcing low-z solution when
 $M_* > 10^{11.7} M_{\odot}$
at
 $z > 3.5$
(Forrest et al. 2024)

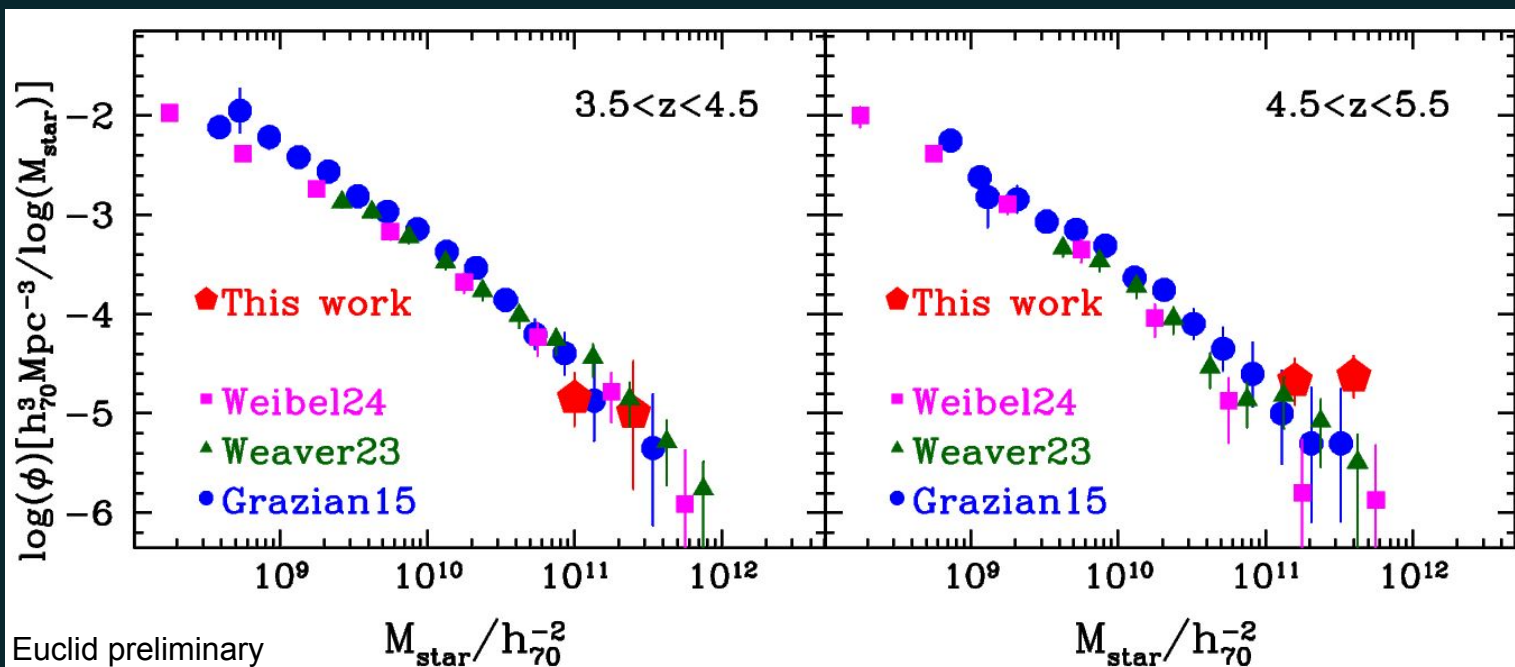


12 candidates, 7 of them with **secondary solution** in Bagpipes analyses

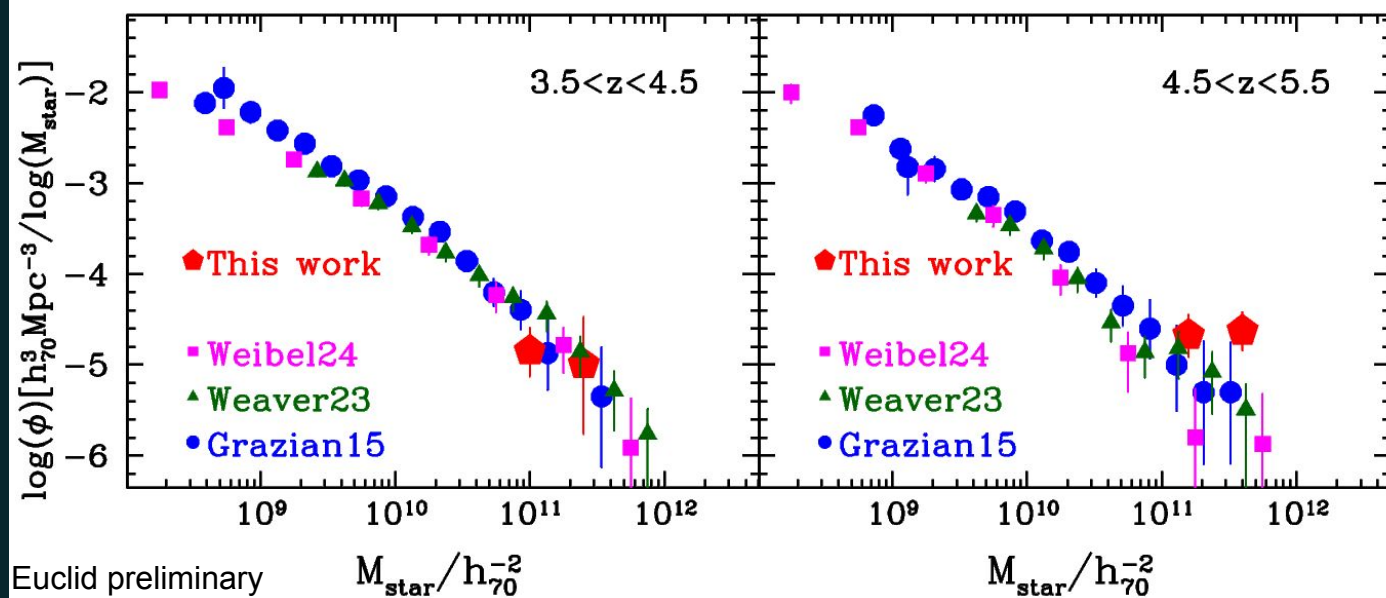
With Bagpipes
 M_* and z



MFs at different
 z bins



Results



- agreement with previous results
- massive end of the MF



What's next?

Just a pilot investigation:
next step **EDF**

~ 40 degrees² overlapping
with Spitzer
(EC: McPartland et al. 2024)



Expected ~ 31.000 **red dusty** sources
Slitless spectroscopy available

Paper in
progress!